

Searches for physics Beyond the Standard Model with the Short Baseline Near Detector

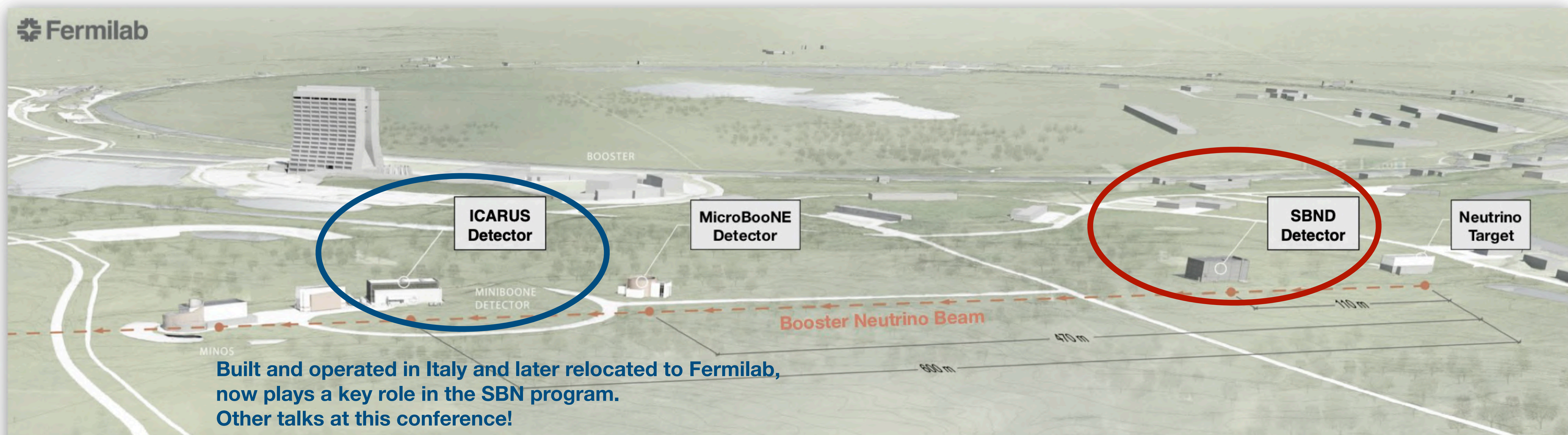


Gaetano Fricano

**PhD Candidate at University of Palermo
On behalf of the SBND Collaboration**

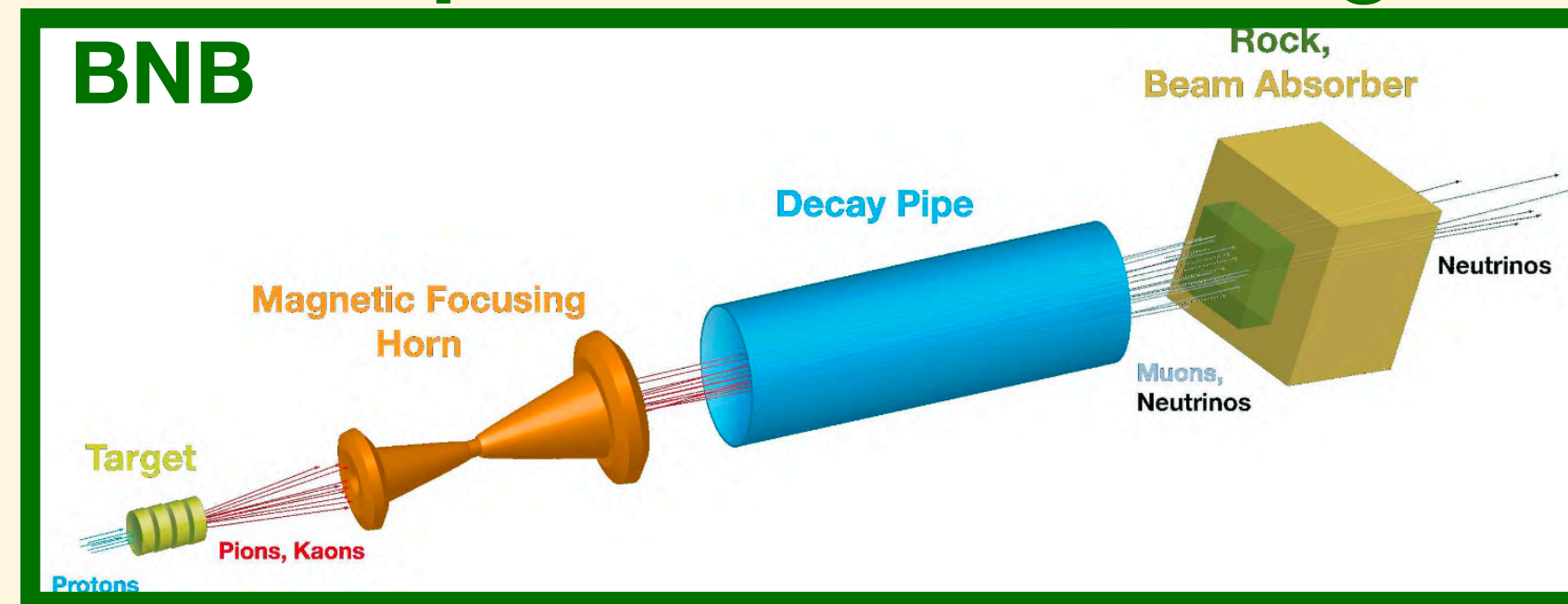


Short Baseline Neutrino Program



Two LArTPC detectors located along the Booster Neutrino Beam (**BNB**) at Fermilab

8 GeV protons on a Be target



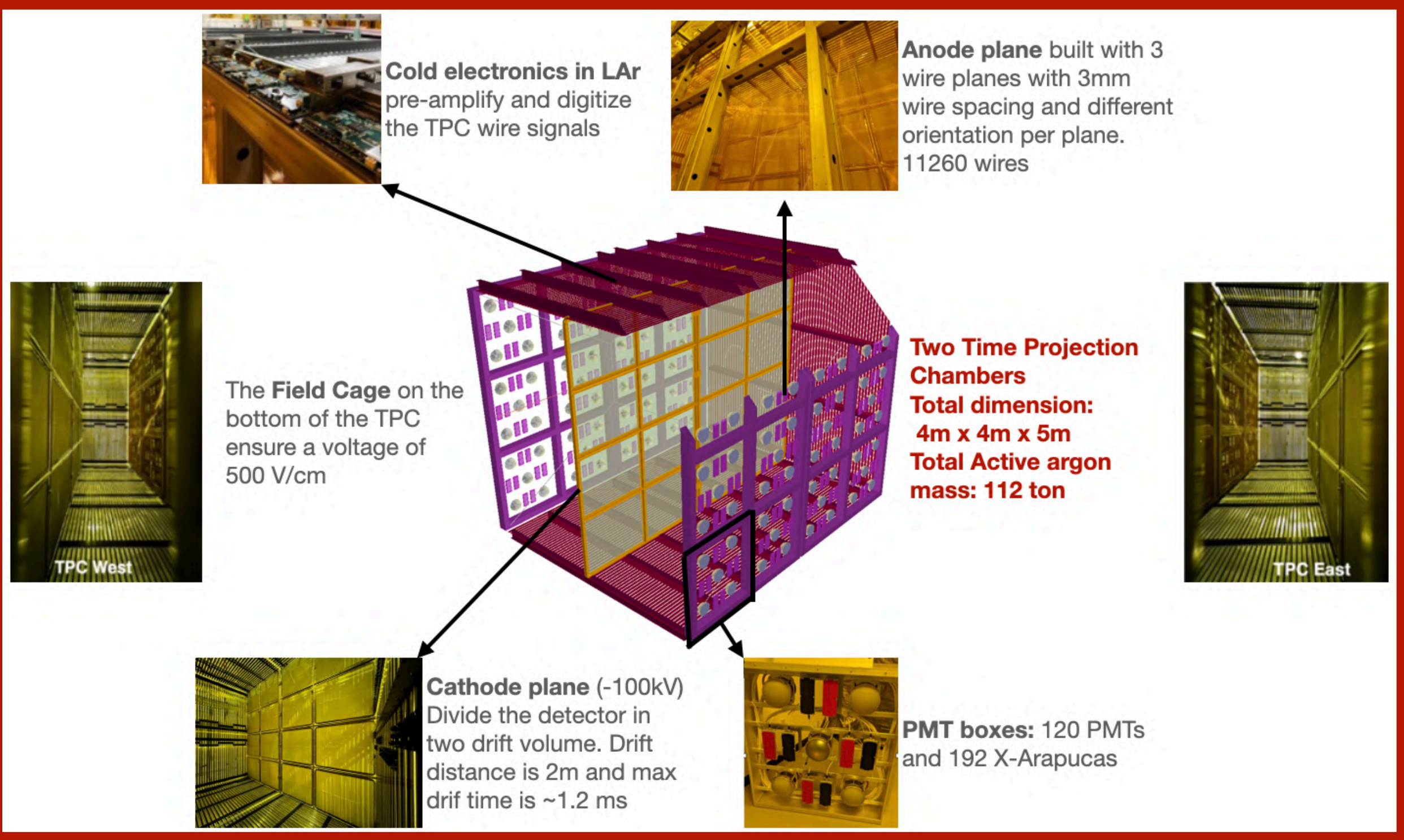
$$\langle E_\nu \rangle \simeq 700 \text{ MeV}$$

SBND

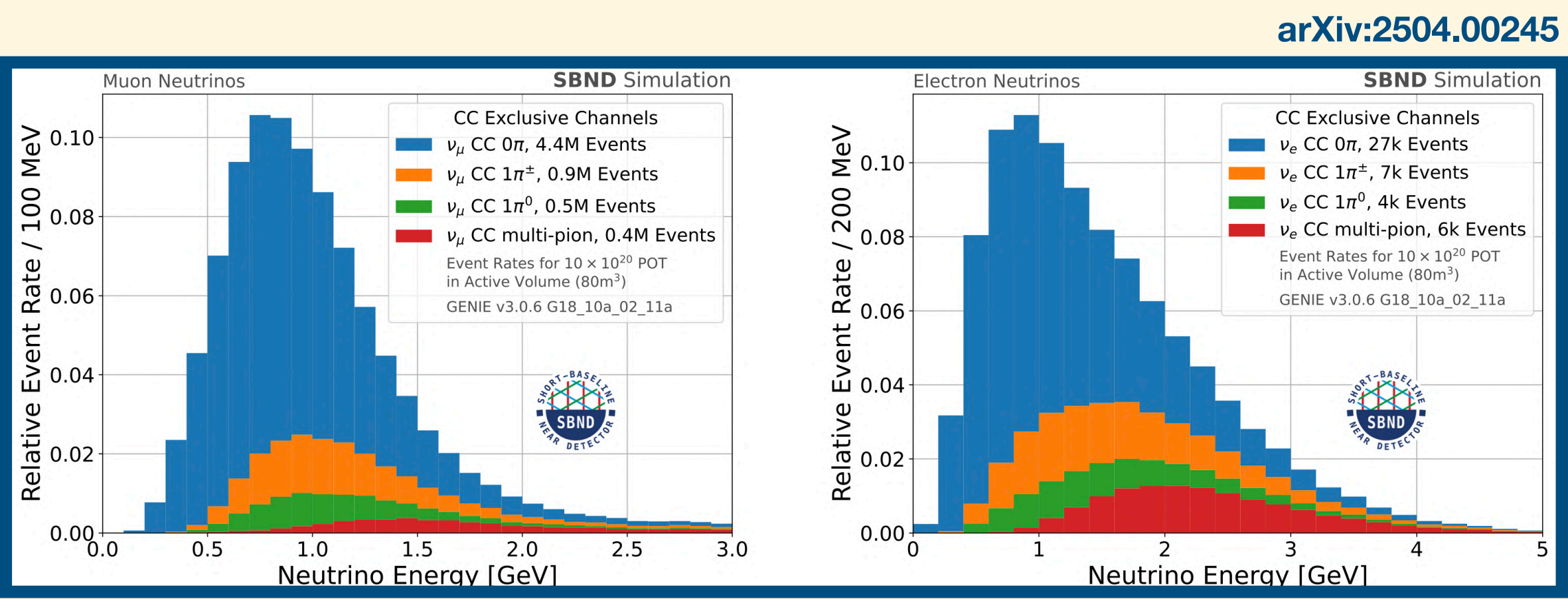


Topic of this talk!

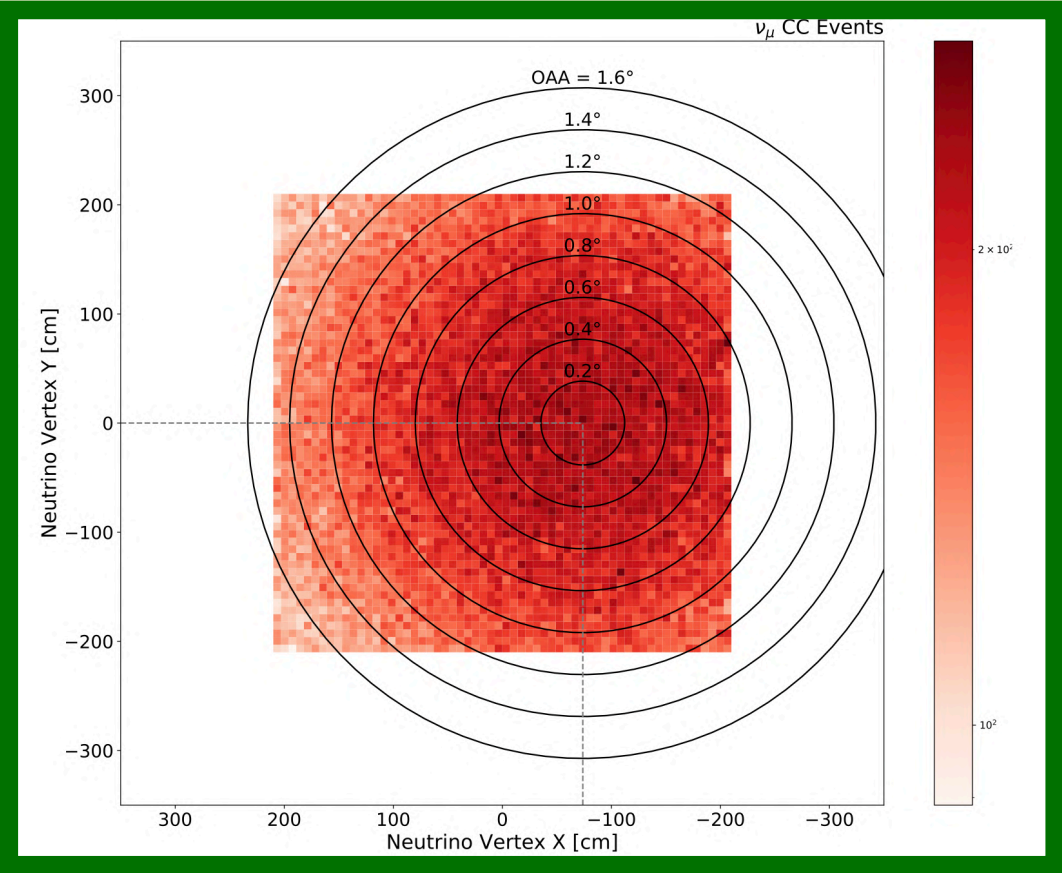
Short Baseline Near Detector



Collects more than **2 million** neutrino interactions per year



6×10^6 CC ν_μ and 45×10^3 CC ν_e in 3 years



SBND PRISM
Flux characterization and background rejection

arXiv:2508.20239

SBND History

December 2022



April 2023



March 2024



July 2024



September 2024



RUN 14737, EVENT 586
July 11, 2024

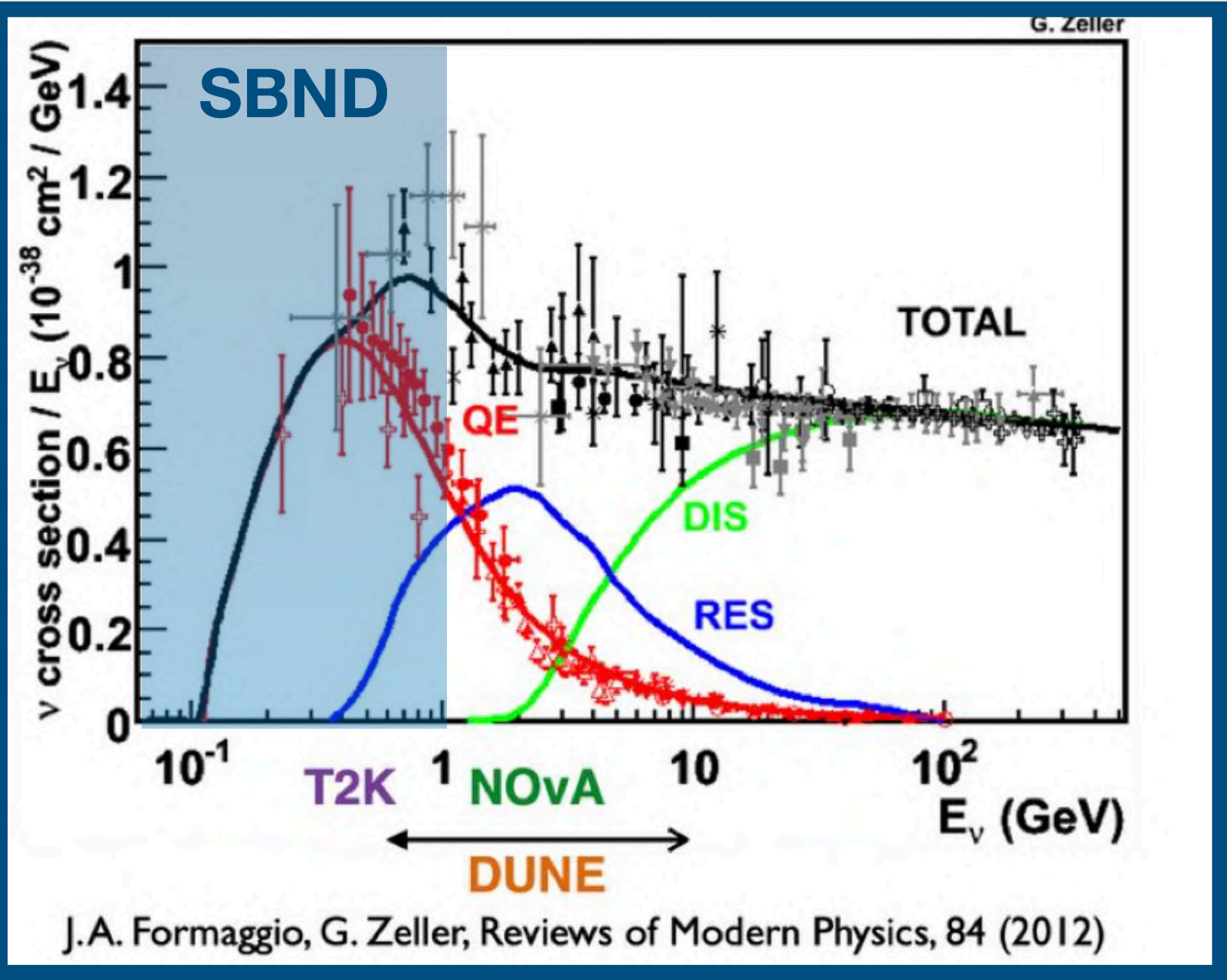
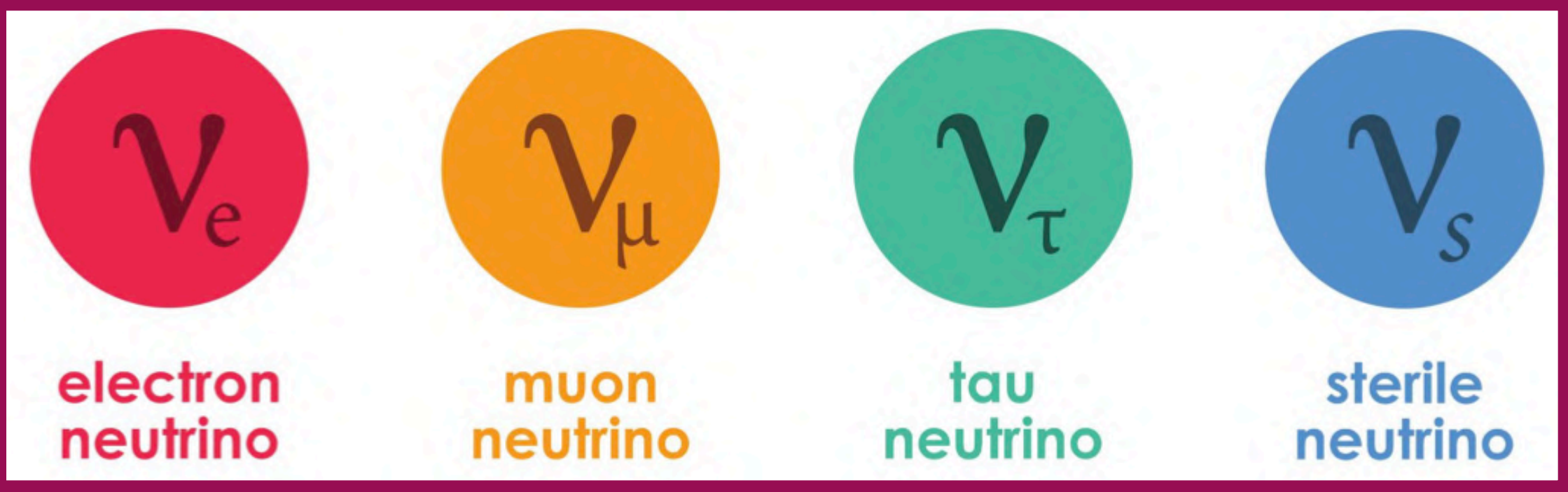
DRIIFT TIME
↑

BEAM DIRECTION → 30 cm

RUN 1: February - July 2025
Already collected the largest neutrino-Argon dataset in the world, with 3 million interactions (2.8×10^{20} POT). Two more years of data taking

SBND Physics Goals

Search for eV Sterile Neutrino with ICARUS



Measure
neutrino cross
sections with
LAr

Beyond the Standard Model Physics

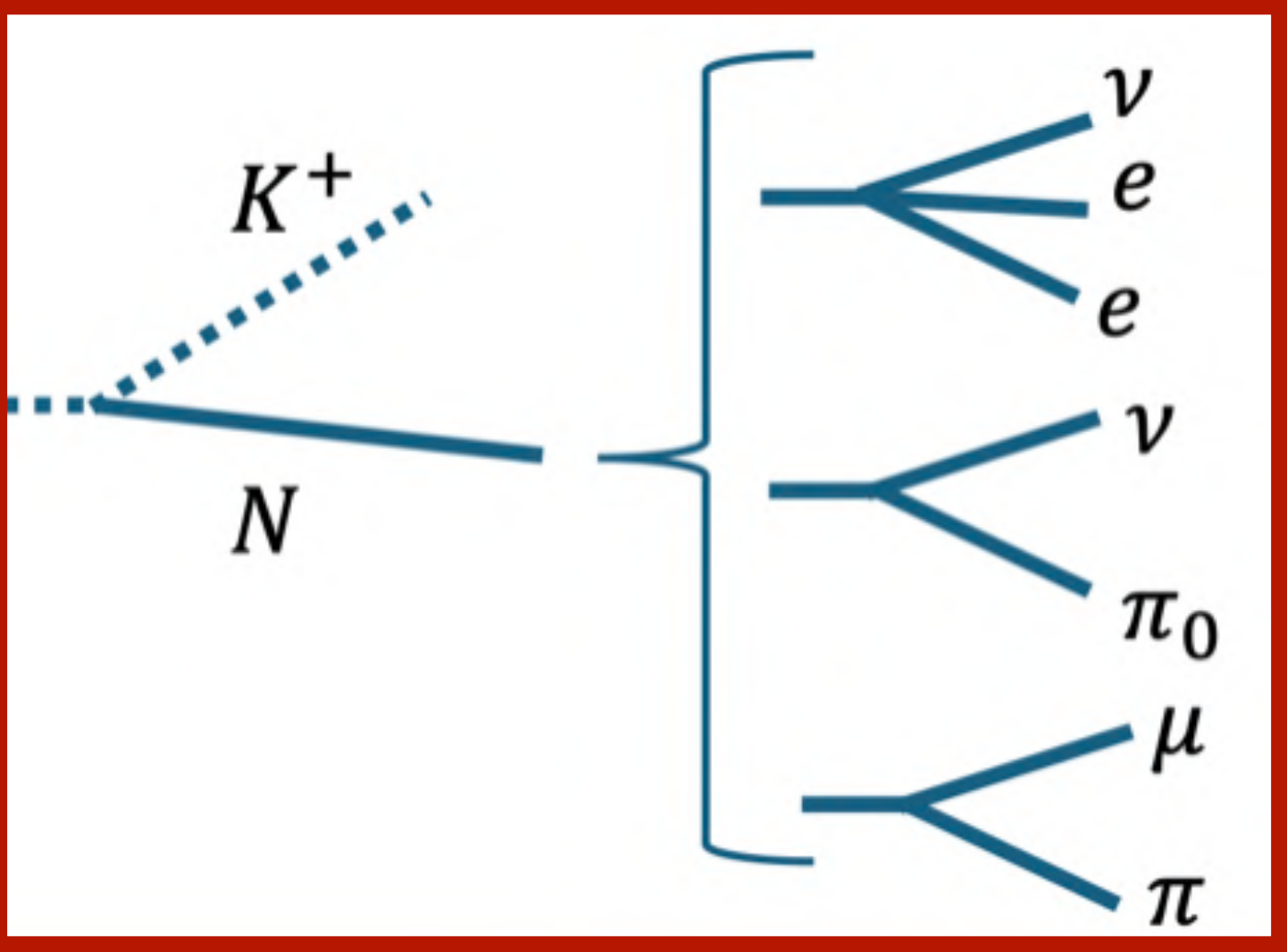
Given its proximity to the target and its excellent spatial, temporal, and energy resolution, **SBND** provides ideal conditions to search for new states produced in the BNB that decay inside its volume, thereby opening the door to dark sector investigations.

Topic of this talk!

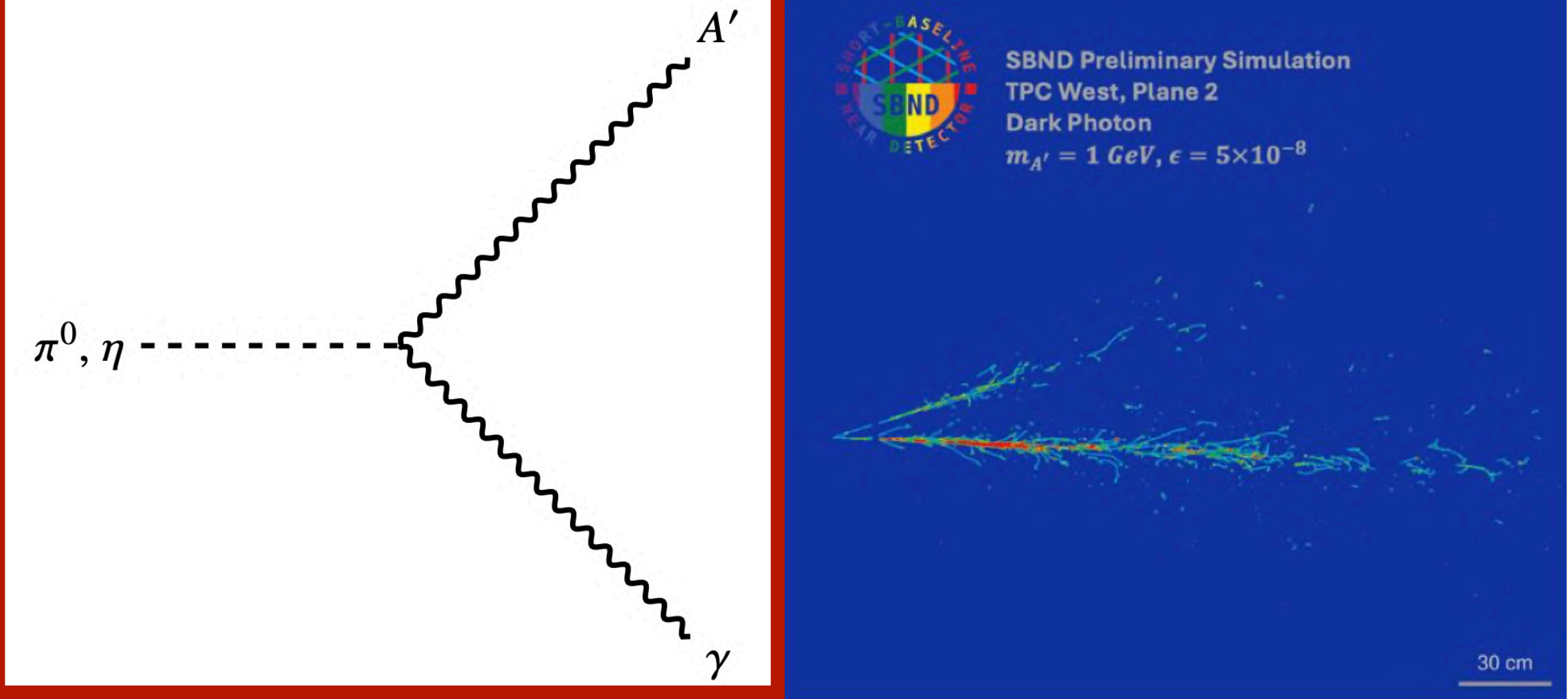


Some of the BSM Models in SBND

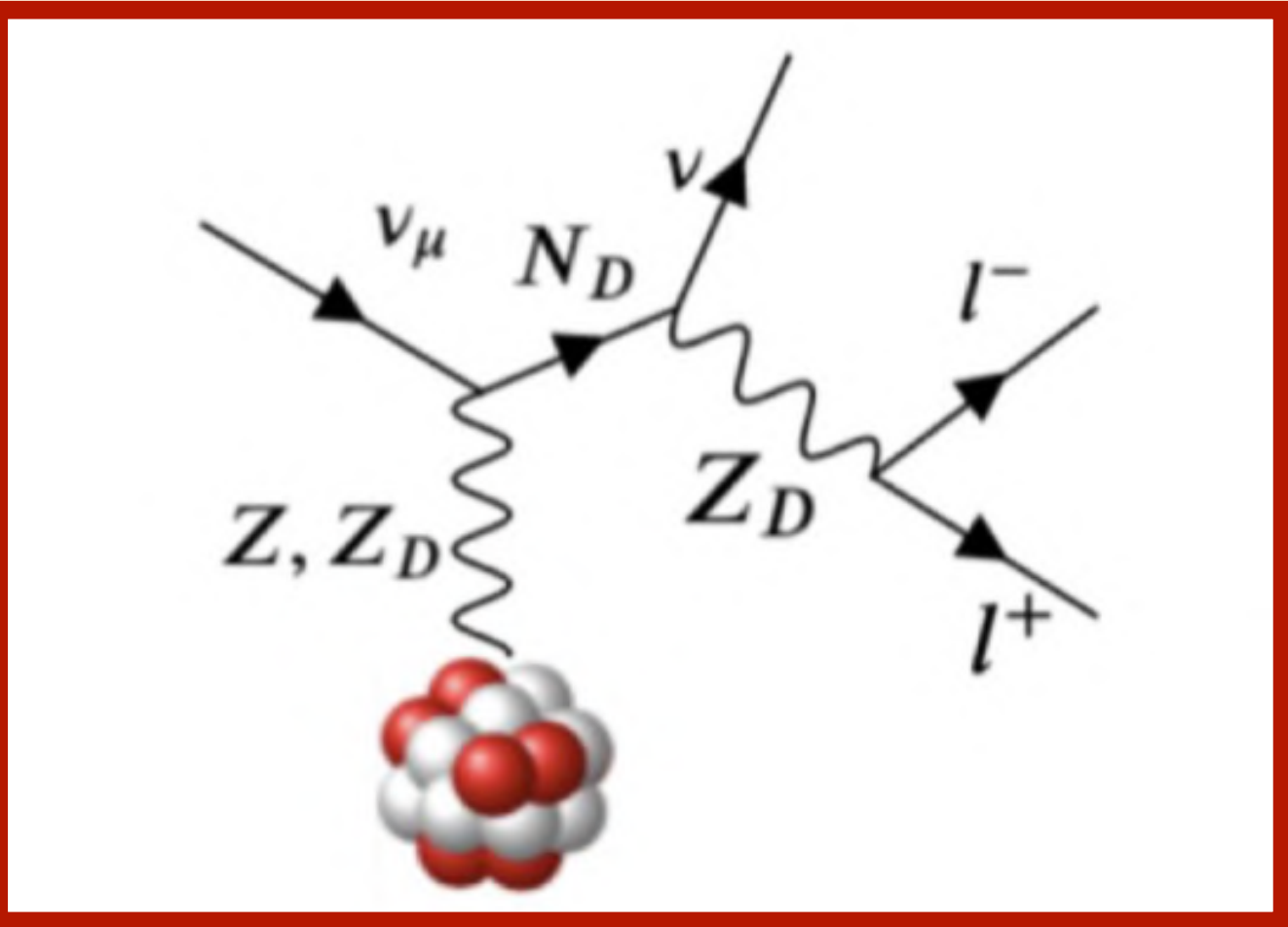
Heavy Neutral Lepton



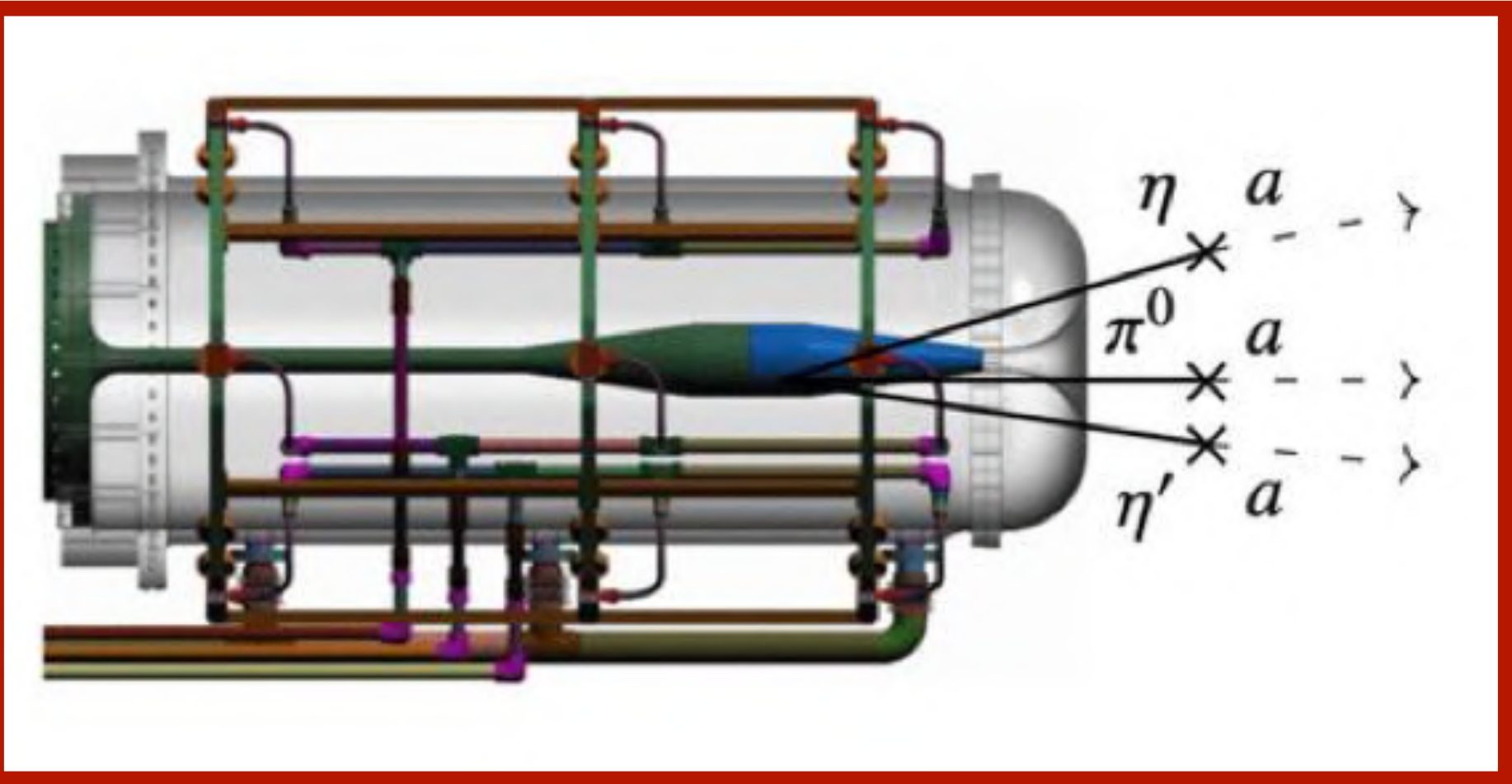
Dark Photons



Generic Long-Lived Particles



Heavy QCD Axions



...To be continued!

How to search for BSM Models in SBND

Kinematic analyses

Energy, momentum, invariant mass, and angular distributions — are essential to separate BSM signals from background

Background

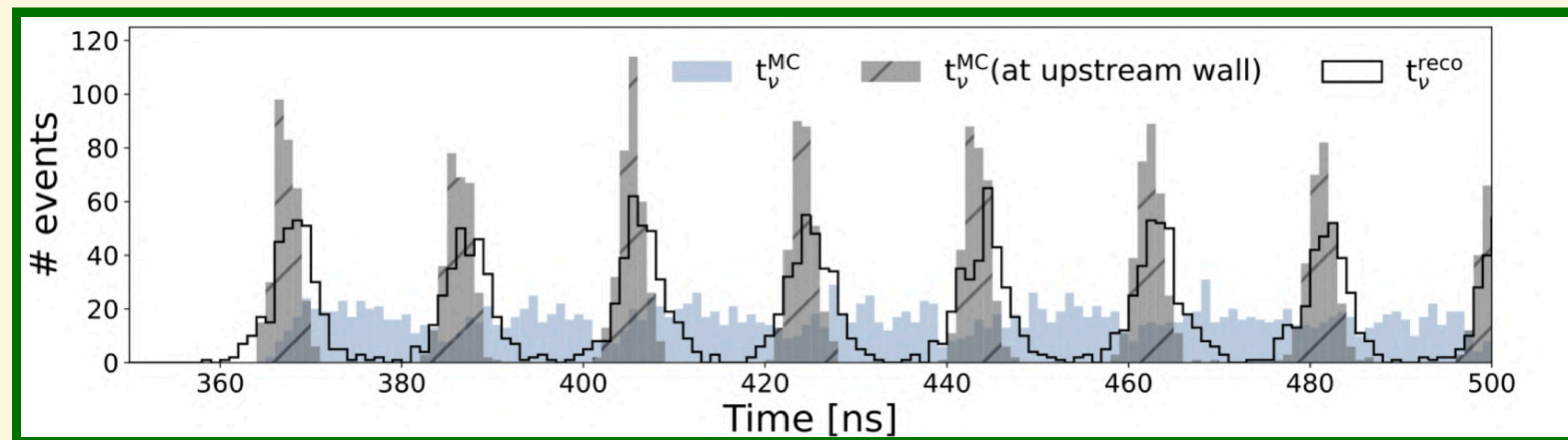
Beam-induced $NC\pi^0$ events and generic cosmic-ray induced ones

Timing analyses

Protons arrive in ~2 ns bunches spaced by ~19 ns.

The PDS provides O(ns) timing resolution

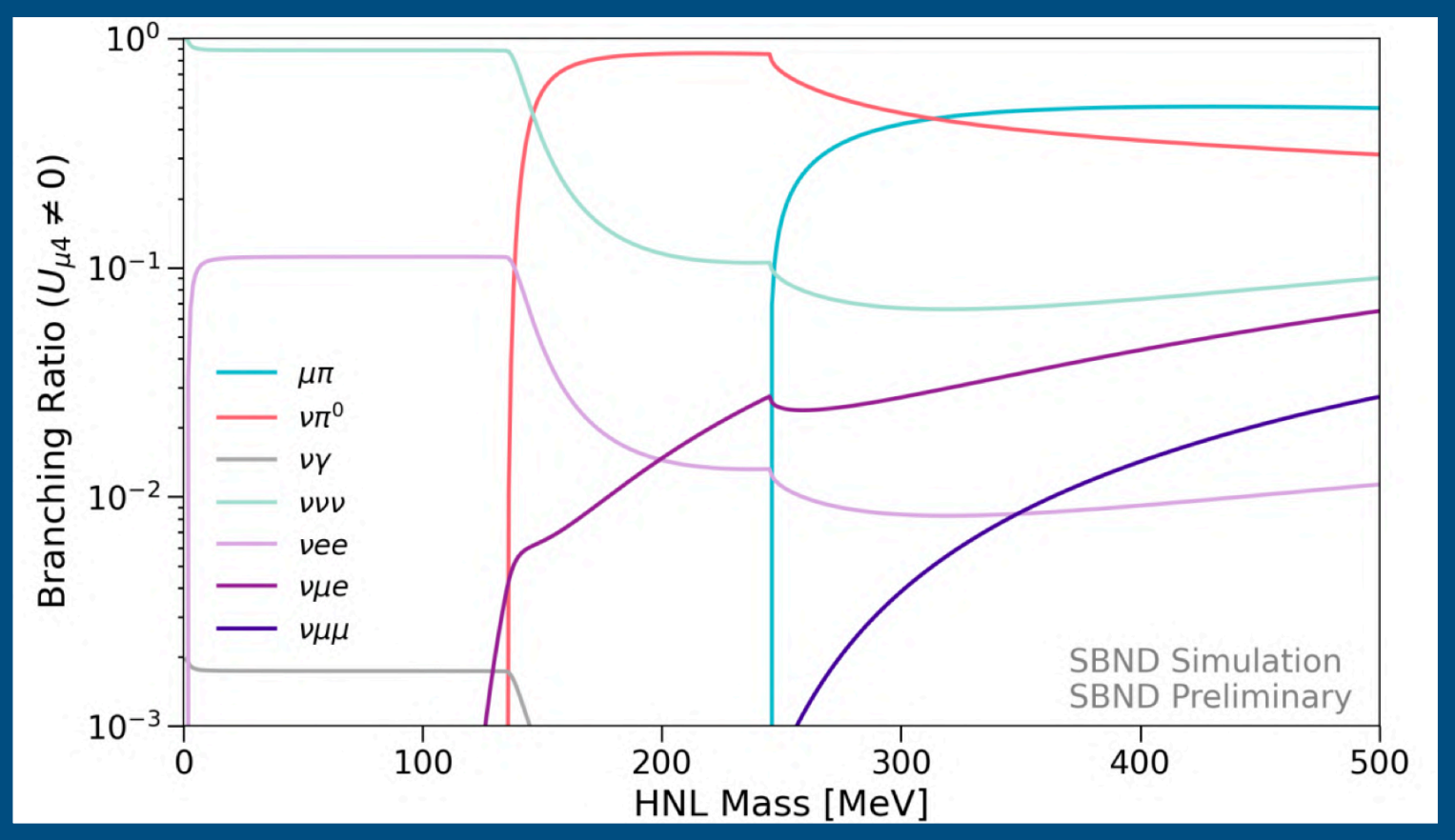
BSM decays may be outside neutrino bunches



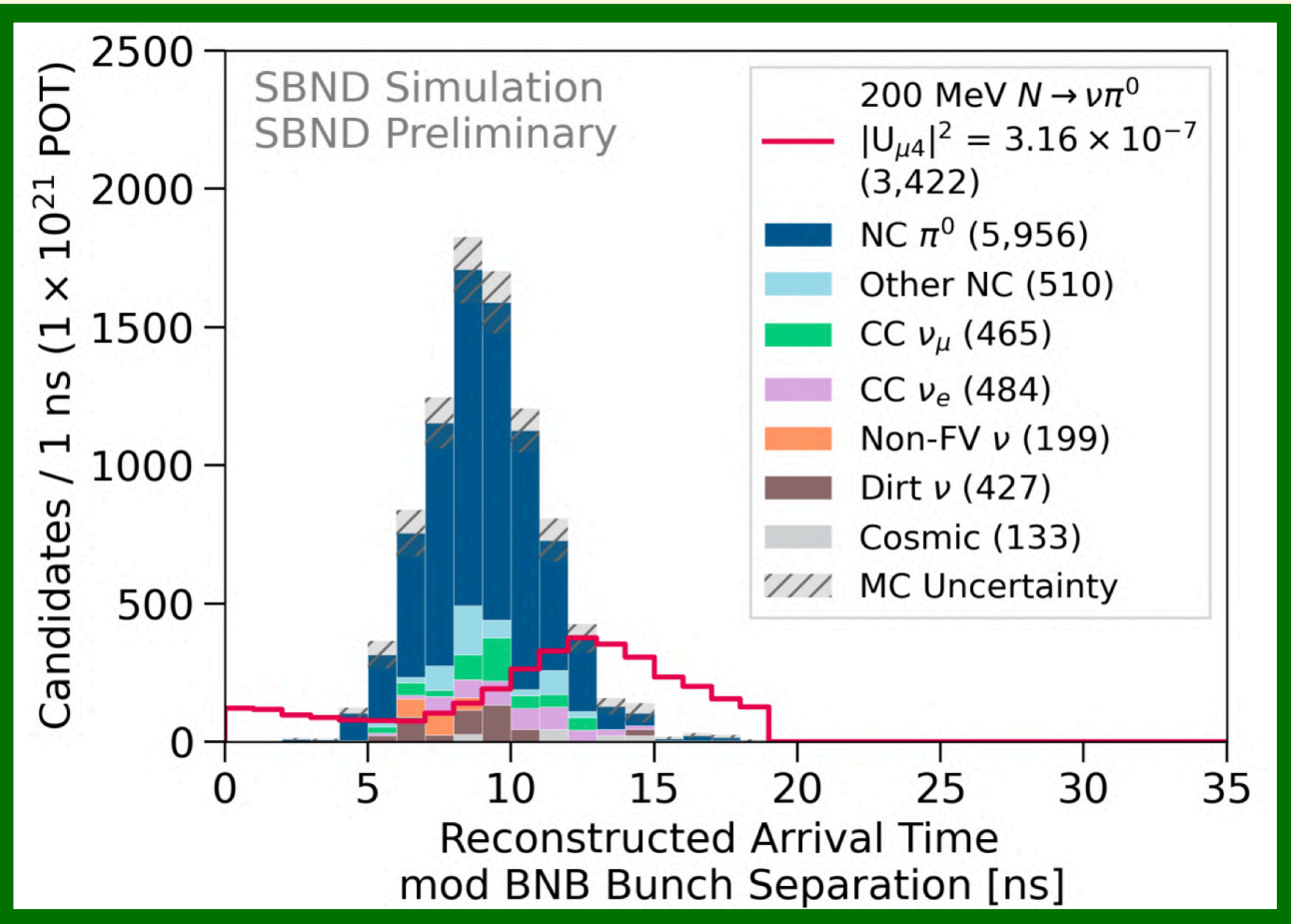
Heavy Neutral Lepton

Proposed as right-handed fermion coupled to the standard neutrinos via an extended PMNS matrix. Assumed as Majorana one for this analysis

Production and decay

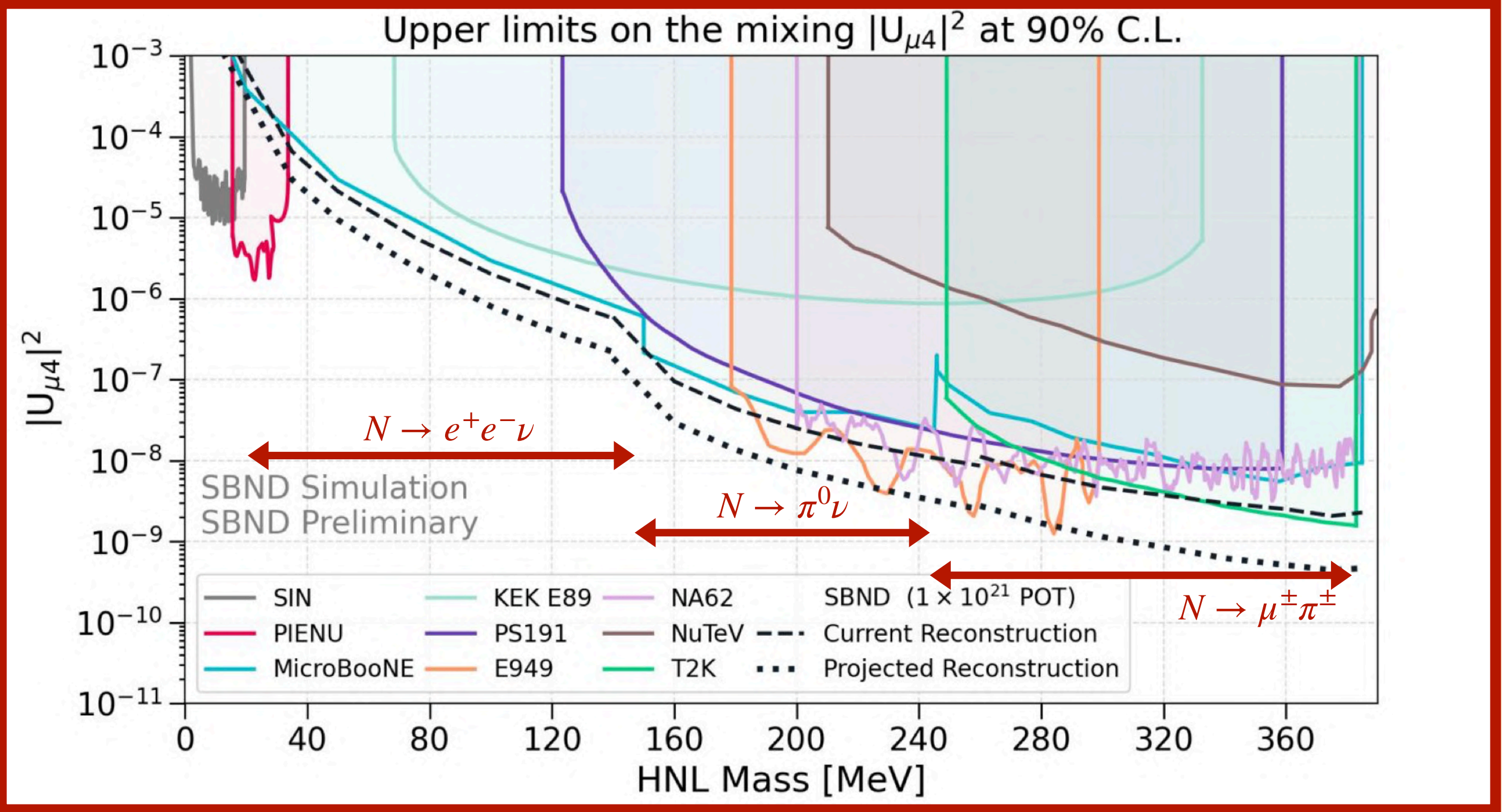


MeVPrtl
simulation



Exploiting time
delays relative to
neutrinos
(arrival time)

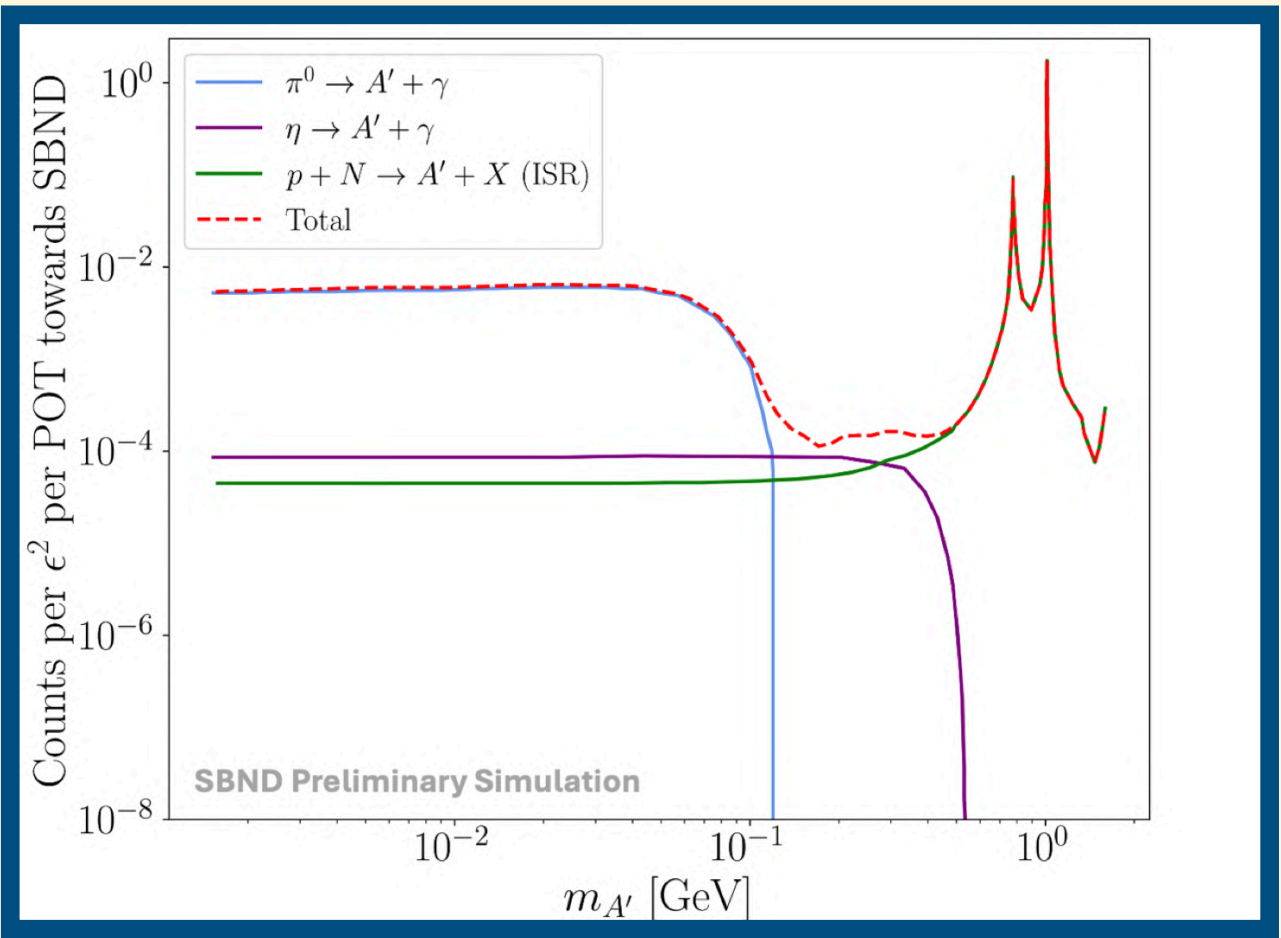
Sensitivity



Ability to place competitive, world-leading
bounds on the HNL parameter space

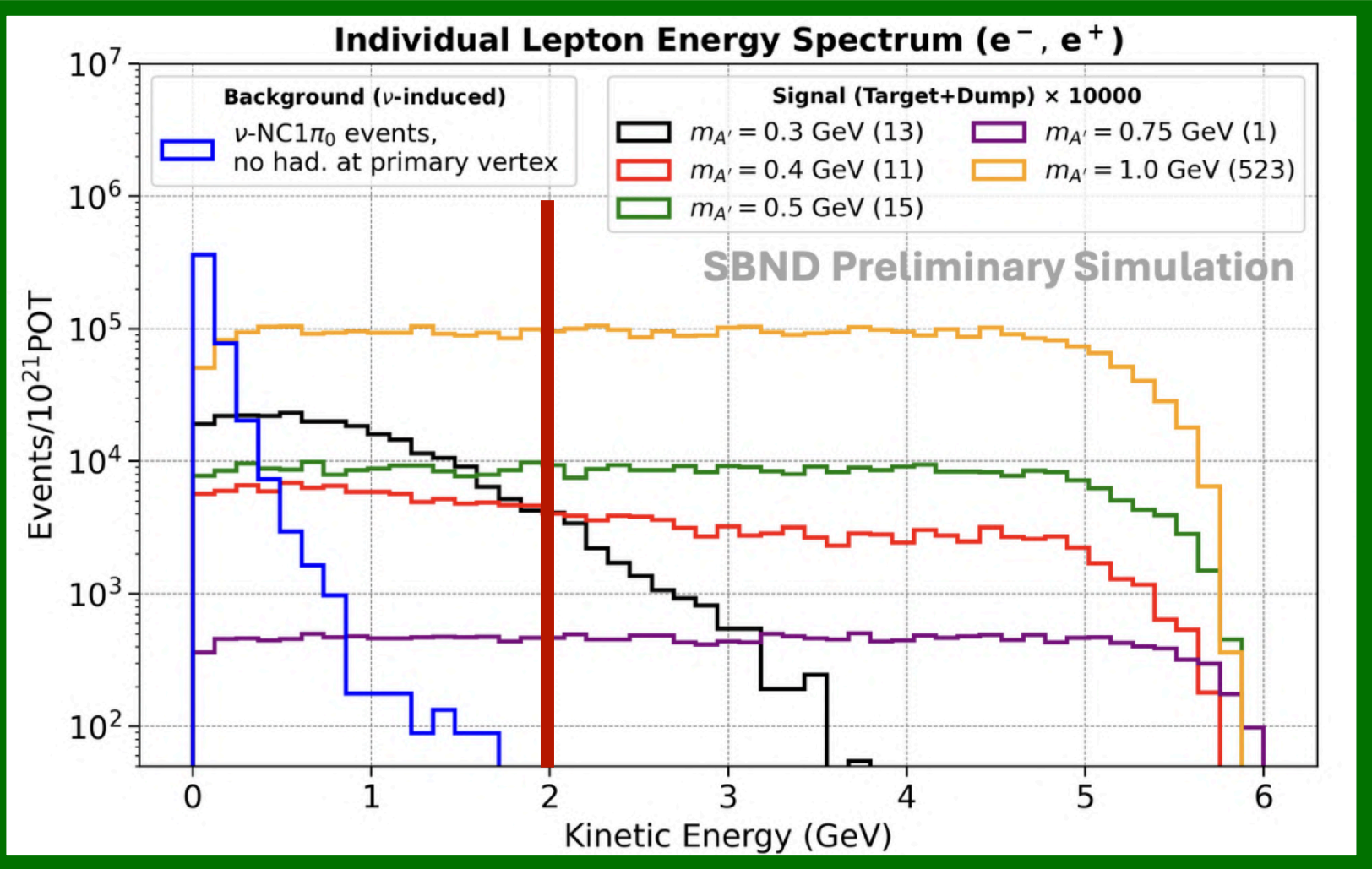
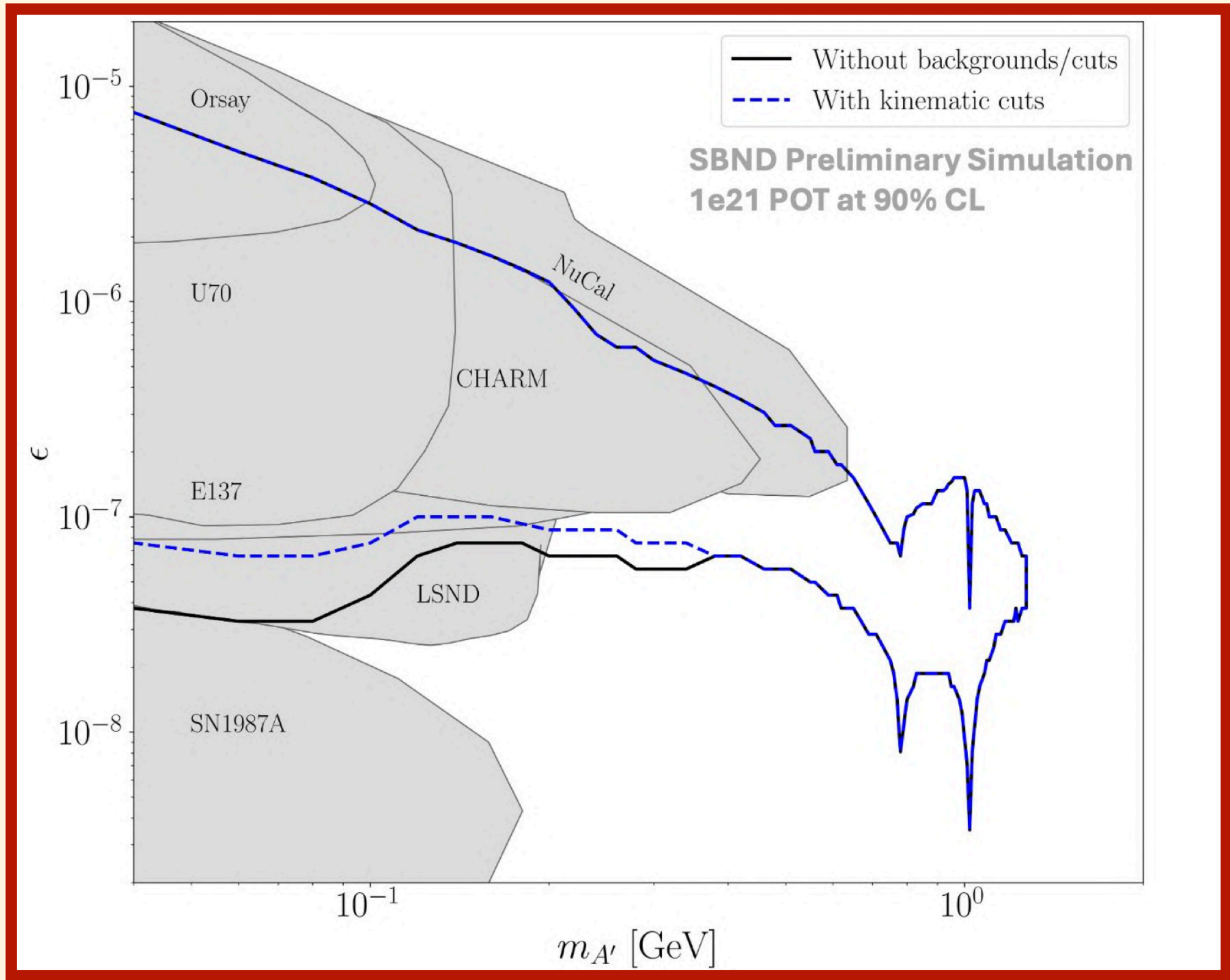
Dark Photon

U(1) gauge boson coupling to the SM via ϵ , kinetic mixing $\mathcal{L} \supset \frac{m_{A'}^2}{2} A'_\mu A'^\mu + e\epsilon \sum_f q_f \bar{f} \gamma^\mu f A'_\mu$



Production

Sensitivity



The higher kinetic energy of electromagnetic showers induced by Dark Photons can be exploited by applying a 2 GeV cut

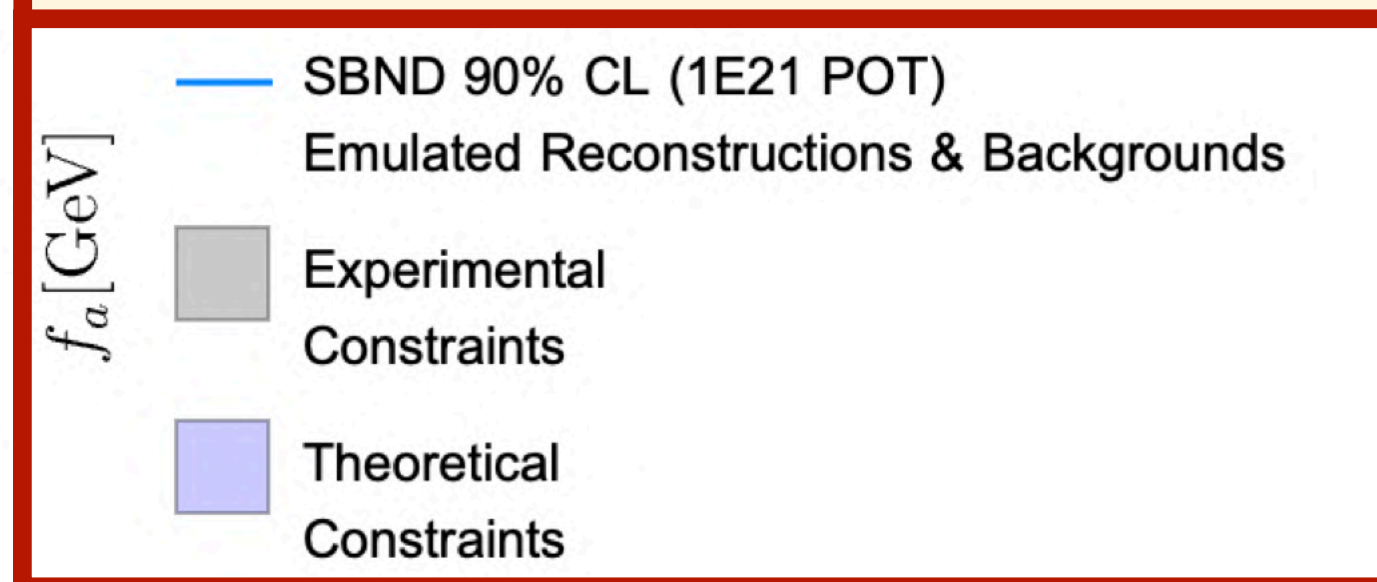
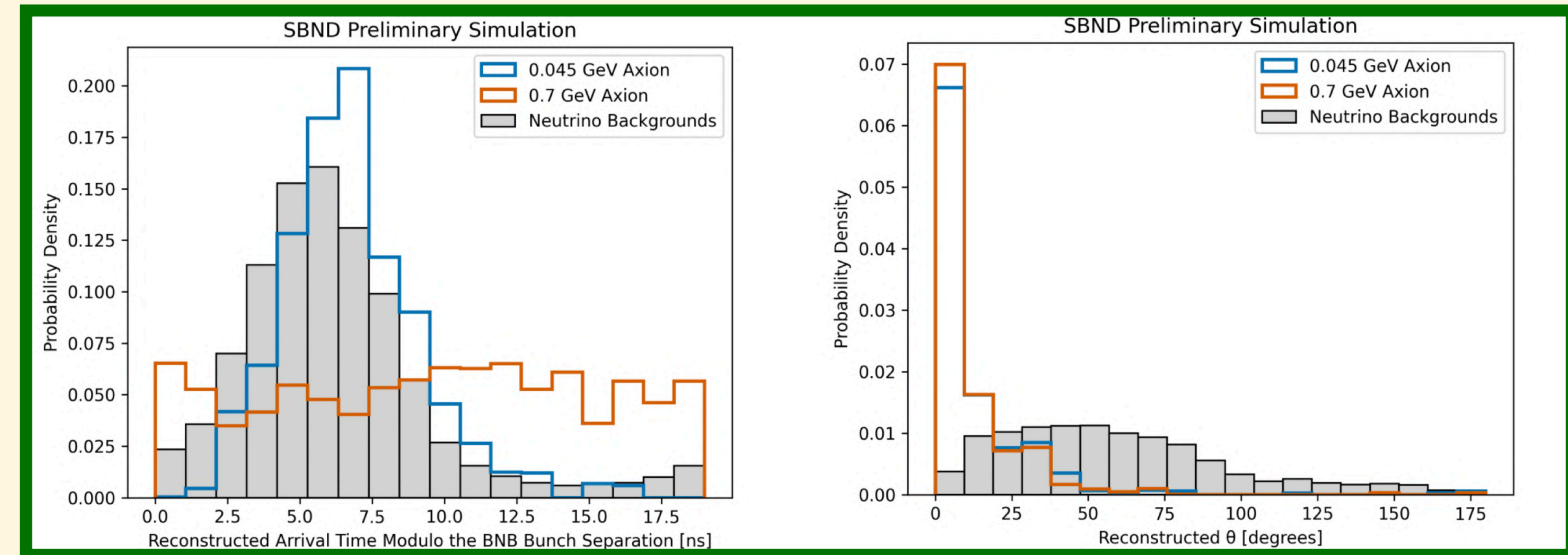
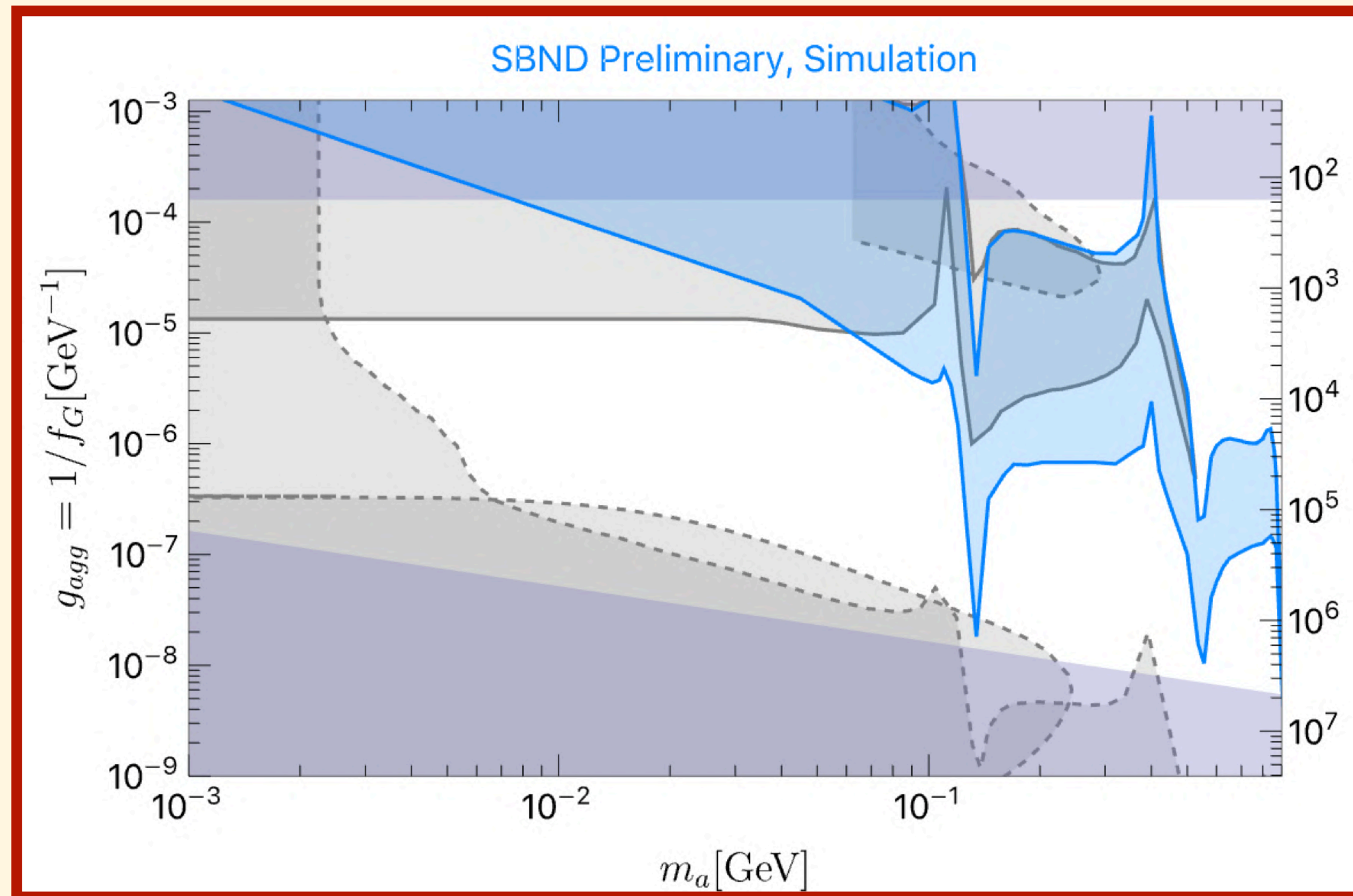
Ability to place world-leading bounds on Dark Photon parameter space

Heavy QCD Axion

Proposed solution to the Strong CP problem $\mathcal{L}_{eff} = \frac{a}{8\pi f_a} (c_3 \alpha_3 G\tilde{G} + c_2 \alpha_2 W\tilde{W} + c_1 \alpha_1 B\tilde{B})$

Produced by neutral mesons in the BNB
Same neutral mesons simulation as Dark photon

Exploiting time delays and angular distribution

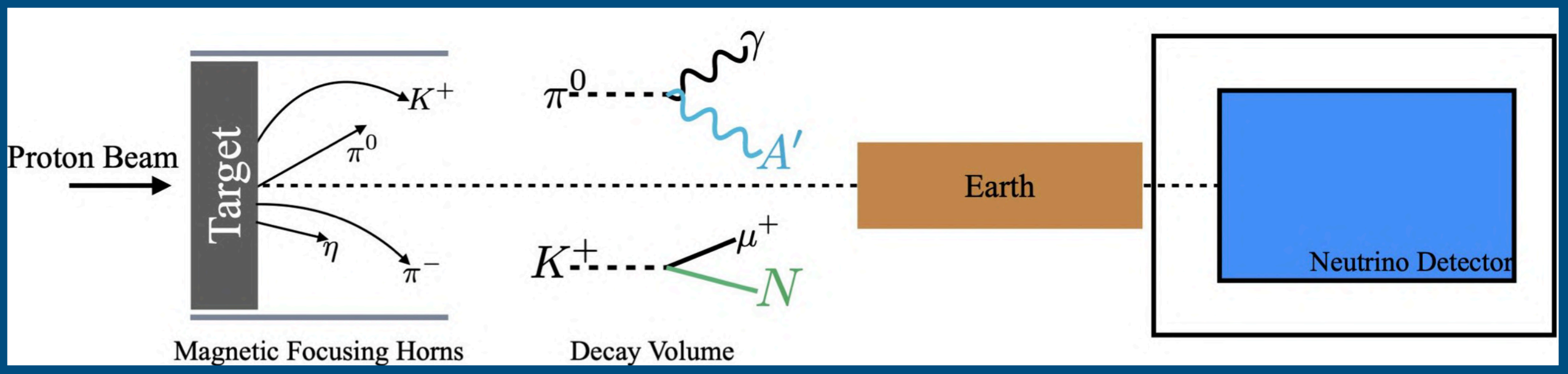


Ability to place world-leading bounds on Heavy QCD axions parameter space

Generic Long-Lived Particles

Model-independent search for exotic particles produced in the BNB and decaying in SBND

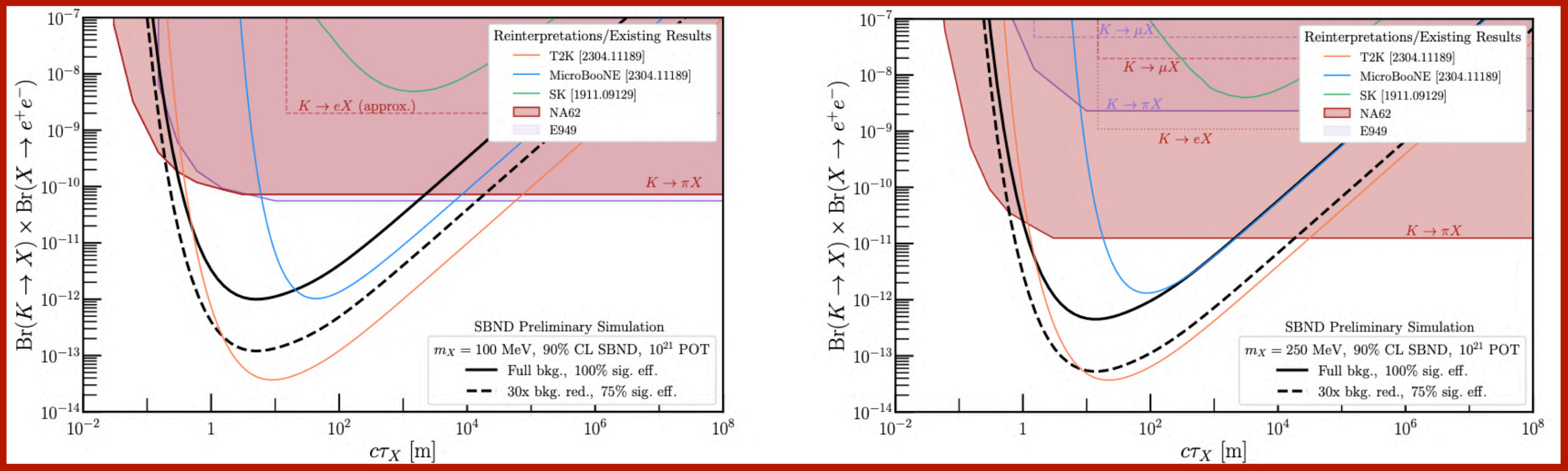
Production



For this analysis we consider the production channels

$$K^+ \rightarrow \pi^+ + X$$
$$K^+ \rightarrow \mu^+ + X$$

Sensitivity



A model-independent search also has the potential to set world-leading limits

Summary and Outlook

- **SBND has capability to probe many BSM models — a publication is under preparation!**
 - **LArTPC and PDS provide excellent spatial and timing resolution**
 - **Timing and kinematic analyses offer the potential for world-leading searches**
- **Ongoing SBND data analyses are expected to produce exciting new insights — stay tuned!**

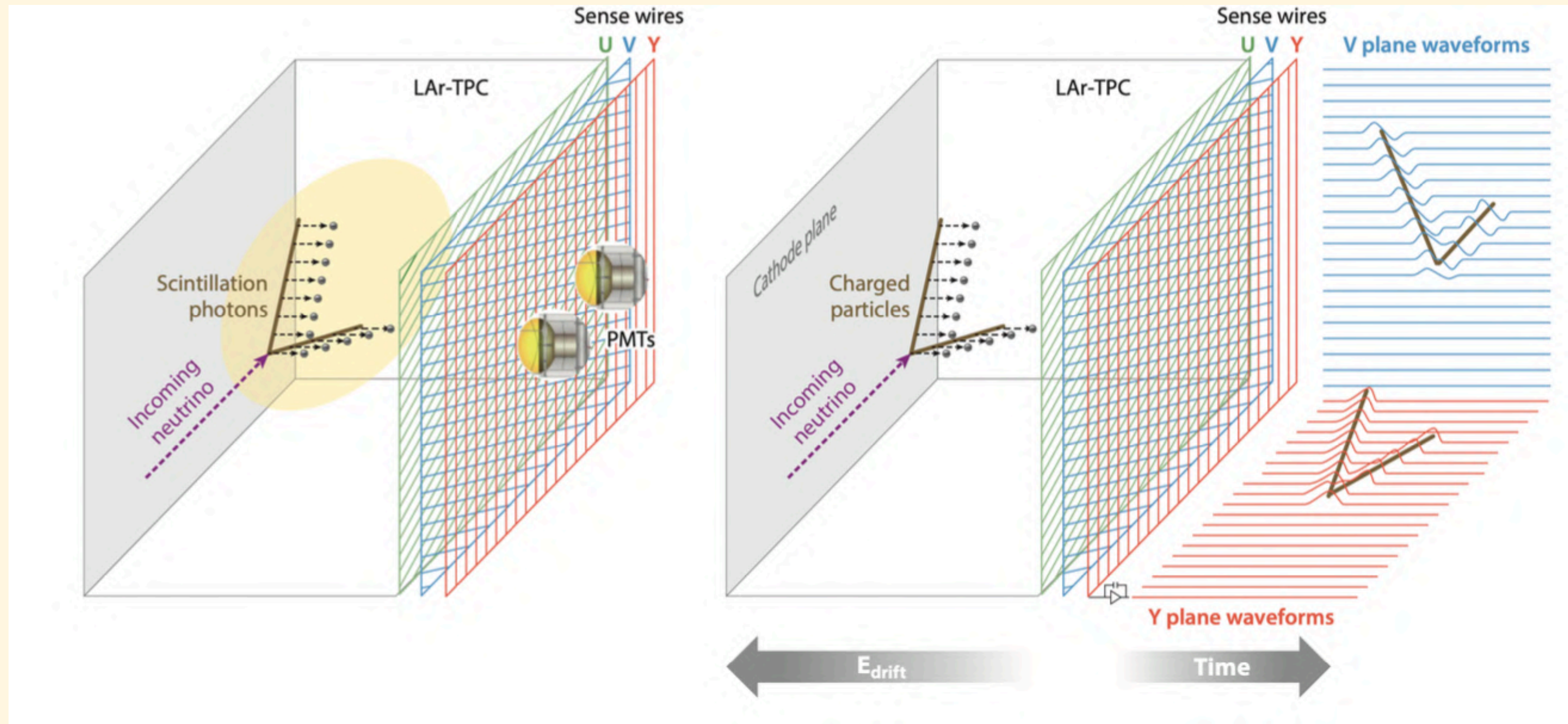


SBND Collaboration Meeting at Sheffield University, UK June 2025

Thank you!

Backup Slides

Liquid Argon Time Projection Chamber



Charged particles produce ionization electrons and scintillation light (event time)

Ionization charges drift in a uniform electric field towards the readout wire planes.
Electron drift time $\sim$ ms

Digitized signals from the wire are collected [time of the wire pulse gives the drift coordinate of the track and the amplitude gives the deposited charge]

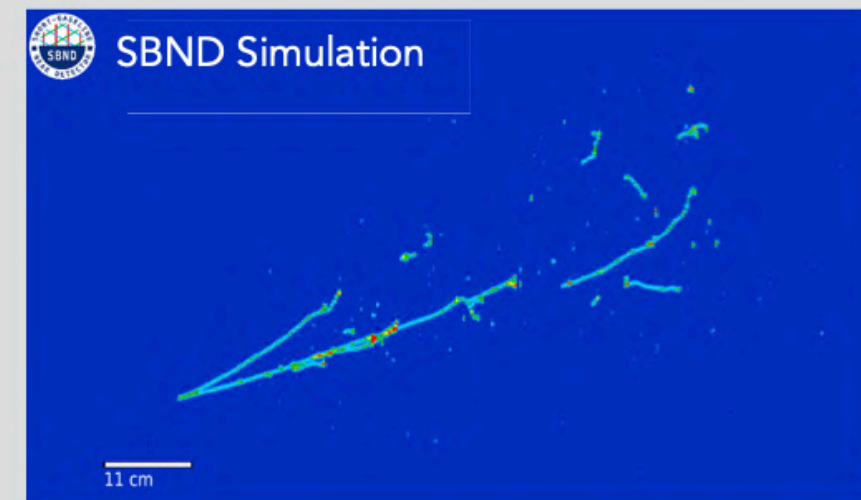
MIP at 500 V/cm:
 $\sim 60,000$ e/cm
 $\sim 50,000$ photons/cm

Other BSM models and event display

Alternative explanations
to the MiniBooNE excess
and other BSM scenarios

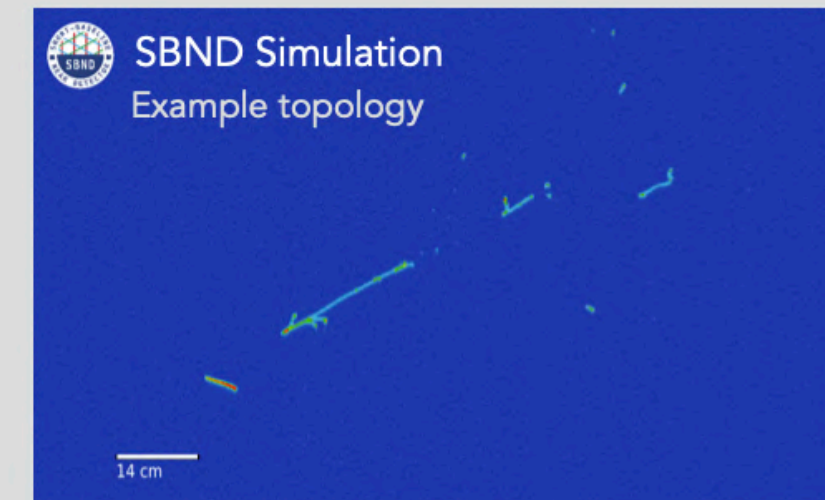
Not an exhaustive list

Dark Neutrinos



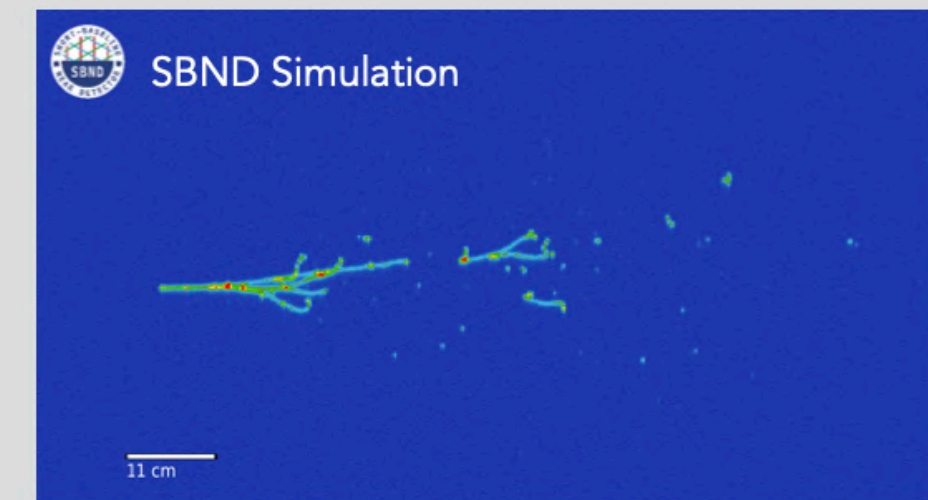
e^+e^- pair w/ or w/o
hadronic activity

Transition Magnetic Moment



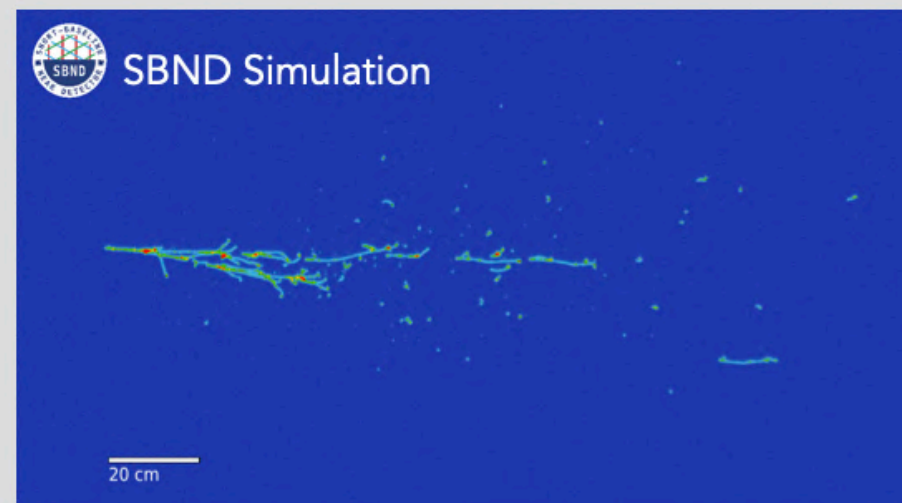
photon shower and
hadronic activity

Axion-like Particles



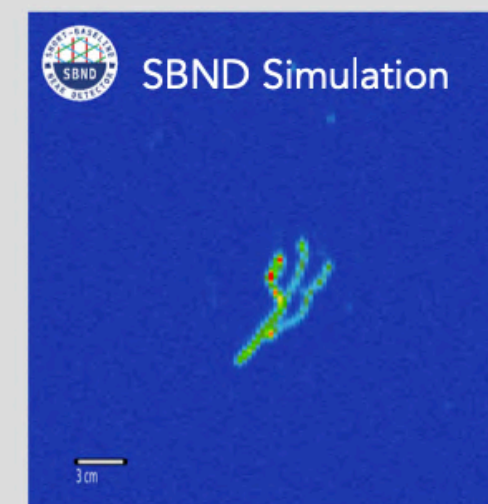
high-energy
 e^+e^- , $\mu^+\mu^-$

Heavy Neutral Leptons



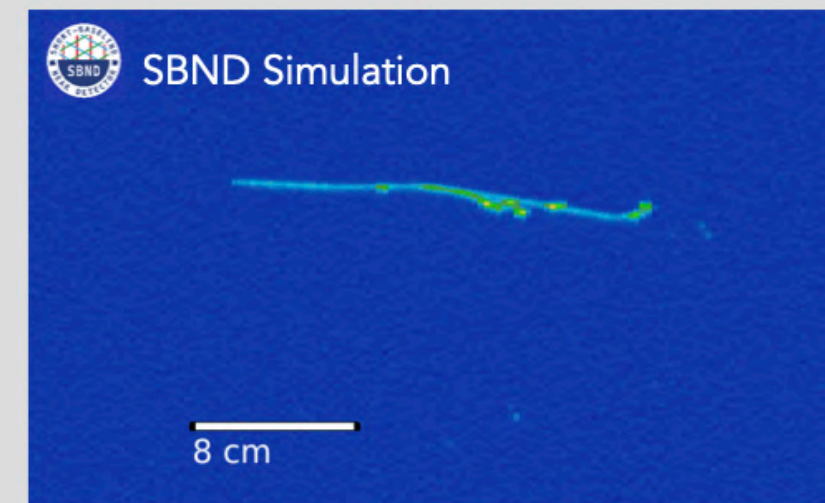
e^+e^- , $\mu^+\mu^-$, $\mu\pi$

Higgs Portal Scalar



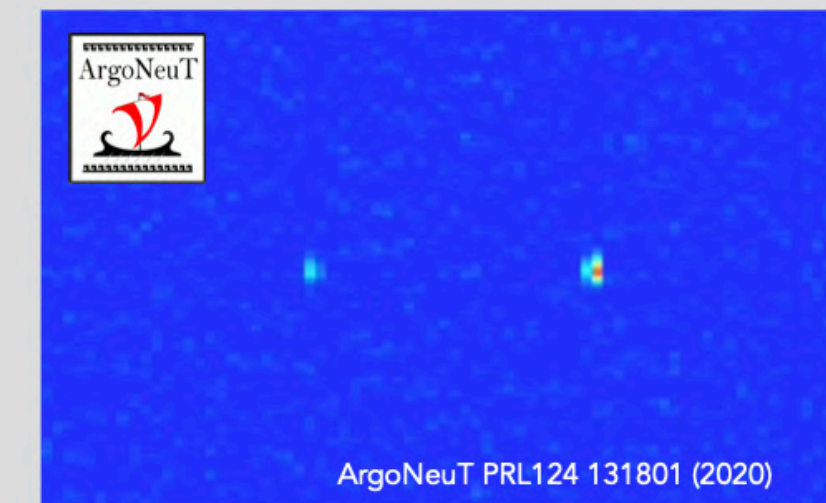
e^+e^- , $\mu^+\mu^-$, no
hadronic activity

Light Dark Matter



electron scattering

Millicharged Particles



blips/faint tracks

PRISM

